

MASEVICH, A.G.

Evolution of early-type stars. Astron. zhur. 32 no.6:498-502  
M-D '55. (MIRA 9:2)

1. Gosudarstvennyy astronomicheskiy institut imeni P.K. Shtern-  
berga. (Stars)

*Mosovich, A. G.*  
 USSR/Astronomy - Variable stars  
 Item 1/1 Feb. 25 - 1/31  
 Author : Mosovich, A. G.  
 Title : ~~Variable stars and their role in Cosmogony~~  
 Publication : ~~Trudy Akad. Nauk SSSR, Ser. Astron., 1975, No. 1, p. 1-10~~  
 Abstract : An account is given of the findings of the Fourth Cosmogony Conference held in Moscow the 26 - 27 October 1974. Reports were read on observations made of pulsating stars and particular attention was given to new stars which flare up and increase their brightness by 10,000 times. An analysis of the data from observations made is said to make possible the drawing of conclusions about their chemical composition, their mass, the amount of matter thrown off and the connection of these stars with other objects.  
 Publishing : ~~Trudy Akad. Nauk SSSR, Ser. Astron., 1975, No. 1, p. 1-10~~  
 Document : ~~Trudy Akad. Nauk SSSR, Ser. Astron., 1975, No. 1, p. 1-10~~

MASEVICH, A.G.

On a visit with India's scientists. Priroda ~~44~~ no.6:57-62 Je '55.  
(India--Scientists--Congresses) (MIRA 8:7)

BAKULIN, P.I., otvetstvennyy redaktor; DUBROVSKIY, K.K., redaktor  
[deceased]; KULAGIN, S.G., redaktor; MASEVICH, A.G., redaktor;  
PARENAGO, P.P., redaktor; RAKHLIN, I.Ye., redaktor; MURASHOVA,  
N.Ya., tekhnicheskiiy redaktor

[Astronomical calender. Yearbook. Variable section for 1957]  
Astronomicheskii kalendar'. Ezhegodnik. Peremennaya chast' 1957.  
Red. kollegiya P.I.Bakulin i dr. Moskva, Gos. izd-vo tekhniko-  
teoret. lit-ry, 1956. 288 p. (Vsesoiuznoe astronomo-geodesiche-  
skoe obshchestvo. no.60) (MLRA 10:3)  
(Astronomy--Yearbooks)

MASEVICH, A.G.

Luminosity function of main sequence stars and its interpretation.  
Astron. zhur. 33 no.2:216-221 Nr-Apr '56. (MLRA 9:8)

1. Gosudarstvennyy astronomicheskiy institut imeni P.K. Shternberga.  
(Stars--Radiation)

MASEVICH, A.G.

Subgiants and their relation to main sequence stars. Astron. zhur. 33  
no.3:330-339 My-Je '56. (MLRA 9:10)

1.Gosudarstvennyy astronomicheskiy institut imeni P.K.Shternberga.  
(Stars--Classification)

MASEVICH, A.G.

Evolution of stars in the Hyades cluster. Astron. zhurn. 33 no.4:576-  
578 J1 - Ag '56. (MLRA 9:11)

1. Gosudarstvennyy astronomicheskiy institut imeni P. K. Shternberga.  
(Stars--Clusters)

MASEVICH, A.G.

Possible courses of continuous evolution of main sequence stars with  
constant and varying masses taking into consideration various. Soob.  
GAISH no.99:3-32 '56. (MLRA 10:3)  
(Stars--Constitution)



BAKULIN, P.I., otvetstvennyy red.; KULAGIN, S.G., red.; ~~MASVICH~~, red.  
PARHAGO, P.P., red.; BAKHLIN, I.Ye., red.; AKHLANOV, S.N., tekhn.red.

[Astronomical calendar; a yearbook. Variable section, 1958]  
Astronomicheskii kalendar'; Ezhegodnik. Peremennaya chast', 1958.  
Red.kollegiya P.I.Bakulina i dr. Moskva, Gos. izd-vo tekhniko-teoret.  
lit-ry, 1957. 303 p. (Vsesoyuznoe astronomo-geodeticheskoe obshche-  
stvo, no.61) (MIRA 11:2)  
(Astronomy--Yearbooks)

MASEVICH, A.G., red.

[Nuclear processes in stars; collected reports read at the Fifth International Colloquium on Astrophysics held in Liege, on September 10, 11, and 12 of 1953] IAdernye protsessy v zvezdakh; sbornik dokladov, pročitannykh na piatom Mezhdunarodnom kollokviume po astrofizike v L'ezhe 10, 11 i 12 sentiabria 1953 goda. Pod red. A.G.Masevich. Moskva, Izd-vo inostr.lit-ry, 1957. 422 p. (MIRA 14:12)

1. Colloque International D'astrophysique. 5th, Liege, 1953.  
(Astrophysics—Congresses)

MASEVICH, A. G.

"Preparation for Visual Observation of Artificial Earth Satellites,"  
a paper presented at the 7th International Astronautical Congress, 6-12  
Oct 1957, Barcelona.

MASEVICH, A.G.

Evolution of stars in open clusters and stellar associations with  
summary in German]. Vor. kosm. 5:56-83 '57. (MLR. 10/81)  
(Stars--Clusters)

MASEVICH, A.G.

Empirical main sequence and its theoretical interpretation (with  
summary in English). Vol. kosm. 5 140-180 '57. (MLBA 10.3)  
(Stars)

MASEVICH, A.G.

MASEVICH, A.G., kandidat fiziko-matematicheskikh nauk.

The physics of planetary nebulae; conference of the Commission on  
Cosmogony. Vest. AN SSSR 27 no.4:119-120 Apr '57. (MLRA 10:5)  
(Nebulae)

AUTHOR:

Masevich, A.G.

30-8-22/37

TITLE:

On Outergalactic Astronomy and Cosmology (Vnegalakticheskaya astronomiya i kosmologiya)

PERIODICAL:

Vestnik Akademii Nauk SSSR, 1957, Vol.27, Nr 8, pp.94-96 (USSR)

ABSTRACT:

The meeting which took place in Moscow from June 5th to June 7th was devoted to the problems of the above mentioned astronomy and cosmology. In his report V.A. Ambartsumyan gave interesting details on this relatively new branch of science. Observations proved the assumption that stellar fogs are not evenly distributed; there are dense parts and accumulations which may be called "light cloud formations". In spite of this, theoretical research is based upon the assumption that the distribution of luminescent cosmic fog is of uniform character. The reviewer further remarked that already now there is no doubt that an expansion of the universe exists. Another reviewer contributed interesting observation results with respect to spiral fogs. The assumption hitherto held that the spiral fog tails are rotating solar systems is said to have been disproved.

Card 1/2

30-8-22/37

On Outergalactic Astronomy and Cosmology

A.L. Zelmanova gave a report on the relativity theory of the anisotropic heterogeneity of the universe. Some reviewers dealt with the problems of the thermodynamics of the universe and a discussion took place on the so-called fluctuation hypothesis. In conclusion a discussion on general problems of cosmology took place.

AVAILABLE: Library of Congress

Card 2/2



Masevich, A. G. 30-11-23/23

AUTHOR: Masevich, A. G.

TITLE: The Problem of Cosmic Gasdynamics.  
An International Conference in the USA.  
(Problemy kosmicheskoy gazodinamiki.  
Mezhdunarodnaya konferentsiya v SSHA.)

PERIODICAL: Vestnik AN SSSR, 1957, Vol. 27, Nr 11, pp. 140-143 (USSR)

ABSTRACT: The physicists' interest in these problems has constantly increased, as the problem of the acceleration of cosmic rays and their lives in the space of the galactic system, as well as the investigation of the formation of interastral magnetic fields is closely connected with the motion of the so-called interastral gases. Representatives of astronomy, physics and mechanics met in Cambridge (Kembridzhe), USA; this was the third international symposium devoted to problems of cosmic aerodynamics. The report by the Dutchman Van der Kholst (observations of the radioemission on the 21 cm wave) caused great interest. G. Vokuler (USA) reported on the observations made in Australia of the spiral structure of the galactic system. O. Vilson (USA) dealt with the new research data regarding the inner kinetics of the planetary nebulae, G. Myunkh (USA) with the internal motions in the nebula of Orion,

Card 1/3

30-11-23/23

## The Problem of Cosmic Gasdynamics.

R. Minkovskiy (USA) reported on the investigation of the group of fiber-like nebulae in the Swan, R. Davis (England) thoroughly examined the physical conditions in the gas-dust clouds on the basis of the most recent results of the observation of radio-radiation. Much attention in reports and discussions was paid to the problem of the dissipation of energy. Kh. Petchek (USA), L. Birman and A. Shlyuter (German Federal Republic - FRG) talked on this topic. Some speakers dealt with the nature of the magnetic field of the spiral extensions of the galactic system. Very great attention was paid by the conference to the problem of the gas-corona and of the formation of the radio-radiation (S.B. Pikel'ner). V.A. Ambartsumyan talked on the genetic connection of young stars with the diffuse environment. By means of observations made he rejected the hitherto existing assumptions with regard to the formation of the stars from an interstral substance. The members of the soviet delegation made themselves acquainted with the institutions and the organization of the optical observations of artificial earth satellites in the USA. The delegation visited the astrophysical observatory in Cambridge (Massachusetts) and a number of other scientific institutions in the USA. Then the

Card 2/3

30-11-23/23

The Problem of Cosmic Gasdynamics.

report deals with the details of the optical observations of the artificail earth satellites in the USA. The delegation showed great interest for the organization and equipment of the Massachusetts Institute of Technology.

AVAILABLE: Library of Congress

Card 3/3

AUTHOR: Masevich, A. G.

499

TITLE: The evolution of stars in the  $\chi$  and h Per cluster.  
(Evolutsiya zvezd v skopleniy  $\chi$  i h perseya.)

PERIODICAL: "Astronomicheskiy Zhurnal" (Journal of Astronomy),  
1957, Vol. 34, No. 2, pp. 176-182 (USSR).

ABSTRACT: The H - R diagrams for the nucleus and the surrounding association of the double cluster of  $\chi$  and h Per show marked differences. Fig. 1 shows the H-R diagram for the nucleus of the above double cluster and Fig. 2 the H - R diagram for the association of the double cluster. (● - stars of the main sequence, ○ - weak supergiants, □ - bright supergiants. From the data of Johnson and Hiltner (1)). The diagram for the nucleus is similar to the usual diagram for a cluster of an early spectral type. The characteristic difference between this diagram and the diagram for the association (Fig. 2) is the presence in the latter of stars of the main sequence of the earliest spectral types (O5 - B0), which lie on the undeviated (primary) upper branch of the main sequence. A further difference is in the interval between the primary branch and the branches deviated from it. The disposition of the supergiants in Fig. 2 is reminiscent

The evolution of stars in the  $\chi$  and  $h$  Per cluster.  
(Cont.)

of the evolutionary curves of gravitationally contracting stars (2). This is further emphasized by the presence in Fig.2 of a few O - stars situated under the main sequence. For massive stars, such as the supergiants considered here, such an evolution takes place very quickly (in any case a time less than  $10^6$  years is required). A comparative study of the H - R diagrams of the double cluster leads to the conclusion that the age of the nucleus and the association is roughly the same, but while in the association the process of star-formation continues, it has ceased altogether in the nucleus. The deviation of the brightest stars in clusters from the mean line of the main sequence is usually taken to be the result of the evolution of these stars when heterogeneity in the chemical composition first begins as a result of the absence of intermixing of matter between the radiant envelope and convective nucleus (3, 4). The so-called "primary" main sequence can be found theoretically in two ways. The first method (evolution of stars occurs at constant mass with no intermixing) was given by Johnson and Hiltner (1 and 5). The primary theoretical main sequence as calculated by them is

The evolution of stars in the  $\chi$  and h Per cluster.<sup>499</sup>  
(Cont.)

shown in Fig.3 (dashed line). The second method was used by the present author and is described in Refs. 3 and 6. The full line of Fig.3 shows the result of these calculations (absorption law: chemical composition corresponding to that of the sun; energy source: the hydrogen-nitrogen cycle). The results of calculation of evolutionary curves (no intermixing; variable mass) are given in the table on p.180. (Nucleus: first line of numbers. Association: last two lines. First column shows the star under consideration). Further details are given in Ref.8. As can be seen the age  $t$  is higher by one order for the nucleus. The above conclusions are in agreement with Oort's theory (11) on the formation of expanding O - associations. It is argued that Minch's paper on the age of early type supergiants leaves out at least one important consideration. A star of the main sequence can only change into a bright giant of an early spectral type if in it intermixing does not take place, i.e. hydrogen burns out only in the convective nucleus which includes about 0.1 of the mass of the star. A star having the mass  $20 M_{\odot}$  will reach this stage in  $\approx 2 \times 10^6$  years. After this,  $\odot$

499

The evolution of stars in the  $\chi$  and h Per cluster.  
(Cont.)

the "peaceful" evolution of the star ends and it either undergoes a catastrophe or, if its development continues with a contracting nucleus, it will become a red giant. In either case the result will not be a supergiant of an early type. In order that the stars considered by Munch (12) should have an age of  $\sim 2 \times 10^7$  years their mass must be of the order of  $200 - 300 M_{\odot}$ . It is suggested that the supergiants considered by Munch have an age  $\sim 10^6$  years, but they were formed not in the galactic plane but at larger latitudes. 3 figures, 1 table, 12 references, 4 of which are Russian.

State Astronomy Institute  
imeni P. K. Shternberg.

Recd. Oct. 15, 1956.

518

**AUTHOR:** Masevich, A.

**TITLE:** A meeting of the Committee for Cosmogony devoted to the development of work on cosmology. (Soveshchaniye komissii po kosmogonii, posvyashchennoye nerspektivam razvitiya rabot po kosmologii).

**PERIODICAL:** "Astronomicheskii Zhurnal" (Journal of Astronomy), 1957, Vol. 34, No. 2, pp. 311-312.

**ABSTRACT:** Representatives of astronomical, physical and philosophical institutions in Moscow and other towns were present. V. A. Ambartsumyan noted that not enough attention was given in the U.S.S.R. to the problems of cosmology while a large number of papers has appeared in this field in other countries. He suggested a series of problems that could be tackled. A. L. El'manov also noted insufficient attention paid in the U.S.S.R. to non-relativistic cosmology. In the exposition of a whole series of cosmological aspects, in particular, those that have ideological implications, it is necessary to exclude simplifications and dogmatism. Possible fields of research (origin of chemical elements, radioastronomy, applications of thermodynamics and statistical physics etc.) were suggested by other members. A conference is to be called in 1957 on the problems of cosmogony. It will be devoted to extragalactic astronomy and cosmology. An approach has been made to the "Uspekhi



A meeting of the Committee for Cosmogony devoted to <sup>518</sup>the development of work on cosmology. (Cont.)

Fizicheskikh Nauk" and "Voprosy Kosmogonii" to publish review articles on extragalactic astronomy and cosmology. Translations of the appropriate foreign books will be carried out.

Recd. Feb. 23, 1957.

MASEVICH, A.G

AUTHOR: Nikol'skiy, G.M. and Masevich, A.G. 33-3-27/32  
 TITLE: Comments by Nikol'skiy on the paper by A.G. Masevich  
 "Luminosity function for stars of the main sequence" and  
 author's reply.

PERIODICAL: "Astronomicheskii Zhurnal" (Journal of Astronomy),  
 1957, Vol.34, No.3, pp. 493-494 (U.S.S.R.)

ABSTRACT: A.G. Masevich (1) considers the important problem of the  
 evolution of stars of the main sequence. The analysis of  
 this problem is carried out by him in the following way:  
 Suppose that stars are formed continuously and  $n$  stars  
 per second enter the main sequence at the point  $M_0$  (absolute  
 stellar magnitude). Evolution takes place in the same way for  
 all stars which move down the main sequence. If  $t$  is the  
 time of evolution of a star from the point  $M_0$  to the point  
 $M$  (see formula 7 in (1))- then:

$$nt = N \int_{M_0}^M \varphi(M) dM$$

where  $\varphi(M)$  is the luminosity function observed at the pres-

Card 1/4

33-3-27/32

Comments by Nikolskiy on the paper by A.G. Masevich  
"Luminosity function for stars of the main sequence" and  
author's reply. (Cont.)

of the second type ( $M_0$  greater than  $M$ ,  $dM$  negative)  
Masevich used the limits of the integral in eq. (1) incorrectly  
and hence his conclusions are invalidated by this mathematical  
error.  $n$  is always positive. There are two Slavic references.

AUTHOR'S REPLY

Nikolskiy's article is based on a misunderstanding. It is  
well known that in all the possible cases of evolution of stars  
at constant mass, the luminosity and the radius of the star  
increase. Since the mass is constant, it follows that, after a  
time, the star ceases to obey the mass-luminosity and the mass-  
radius relations characteristic of the main sequence and hence  
'leaves' the latter. This was considered in detail in (2) and  
(3) and is not, and cannot be, a consequence of eq. (2) in  
Nikolskiy's note. It is surprising that Nikolskiy should quote  
(2) since this work, although printed, has not been issued and  
is still being stored by the publishers. Conversely, the  
result obtained in a discussion of evolution at constant mass  
in (1) without the transposition of the limits of integration  
in (1) is treated (perhaps not very successfully) simply as a

Card 3/4

33-3-27/32

Comments by Nikol'skiy on the paper by A.G. Masevich  
"Luminosity function for stars of the main sequence" and  
author's reply. (Cont.)

mathematical expression of known facts and could be omitted  
without effect on the conclusions that follow.

The choice between the two alternative evolutionary hypo-  
theses using the luminosity function is based in (1) not on the  
sign of  $n$  but on its magnitude ( $n$  is the number of stars  
being formed and is necessarily positive). Nikol'skiy's criti-  
cism is thus rejected and the conclusions reached in (1) .  
There are 4 Slavic references.

AVAILABLE: Library of Congress

Card 4/4

МАСЕВИЧ А.

33-5-6/12

AUTHOR: Ruben, G. and Masevich, A.

TITLE: An Investigation of Evolutionary Sequences of Homogeneous Stellar Models with a Convective Nucleus. (Issledovaniye Evolyutsionnykh Posledovatel'nostey Odnorodnykh Zvezdnykh Modeley s Konvektivnym Yadrom.)

PERIODICAL: Astronomicheskii Zhurnal, 1957, Vol. 34, No. 5, pp. 724-738 (USSR).

ABSTRACT: A detailed calculation using the stellar model with a convective nucleus and the absorption law  $K = K_0 \rho^{0.875} T^{-3.5}$  has been carried out by one of the present authors in Refs. 1 and 2. Possible ways of developing this model in the case of homogeneous and inhomogeneous chemical composition were considered in application to stars of the main sequence. In the present work the possibility of an application of such a model to the problem of structure and evolution of stars of other sequences in Russell's diagram is considered. The authors start with a model having the same chemical composition in both the shell and the convective nucleus. The carbon cyclic reaction is taken as the source of energy according to  $\epsilon = \epsilon_0 XZ \rho T^n$ . The continuous evolution (as a result of gradual transformation of hydrogen into helium) of such a model is considered.

Card 1/4

33-5-6/12

An Investigation of Evolutionary Sequences of Homogeneous Stellar Models with a Convective Nucleus.

in the case of constant and variable mass. Using the notation of References 1 and 2 the relation between the luminosity  $L_1$  and the stellar mass  $M_1$  is written in the form  $L_1 = M_1^\gamma$ . Figure 1 shows the calculated dependence of  $\lg L_1$  on  $\lg R_1$  for different values of  $\gamma$  where  $R_1$  is the relative radius. The relation between these two quantities is linear. An analysis is given of the effect of different parameters on the form of evolutionary curves. Various possible laws of change of mass are considered (different  $\gamma$  in Ref. 5). In each of the models there is a limiting value of  $\gamma$  which depends on the form of the law of formation of energy but is almost independent of the model itself. Within the limits of each possible  $\gamma$  there are certain maximum values of  $M_0$  and  $R_0$  the absolute magnitude of which depends on the accepted model. It is shown that the theoretical curve corresponding to  $\gamma = 3.9$  represents the main sequence quite well. Using other values of  $\gamma$  one obtains evolutionary sequences which do not correspond to real stellar sequences for which the mass is a function

Card 2/4

33-5-6/12

An Investigation of Evolutionary Sequences of Homogeneous Stellar  
Models with a Convective Nucleus.

of both the luminosity and radius. In the case of  $\gamma = 3.9$  both  $M_0$  and  $R_0$  reach their maximum values at the same value of  $Z$ , where  $Z$  is the content of elements heavier than helium. Using results obtained for  $\gamma \neq 3.9$  it is shown that the structure and evolution of a sub-dwarfs can be explained by the present model if one assumes that the amount of heavy elements in them is about 20 times less than in stars of the main sequence. Theoretically possible masses of such stars are comparable with the masses of real sub-dwarfs. On the other hand sub-giants can be explained on this model if one assumes that the amount of heavy elements in this group is four to five times higher than in the stars of the main sequence. This is in agreement with results obtained earlier (Ref. 5). It is pointed out that although it is possible to explain the structure of both sub-dwarfs and sub-giants on the above model using certain assumptions as to the heavy element content relative to the stars of the main sequence it must nevertheless be remembered that the necessary condition in all the calculations is full inter-mixing (same chemical composition in shell and nucleus) which in general may not be observed. There are 11 figures,

Card 3/4

33-5-0/12

An Investigation of Evolutionary Sequences of Homogeneous Stellar Models with a Convective Nucleus.

8 tables, 5 references, all of which are Slavic.

SUBMITTED: April, 12, 1957.

ASSOCIATION: State Astronomical Institute, imeni P.K. Shternberg, Potsdam Astronomical Observatory, German Democratic Republic. (Gos. Astronomicheskii In-t im. P.K. Shternberga, Potsdamskaya Astronomicheskaya Observatoriya, Germanskaya Demokraticeskaya Respublika.)

AVAILABLE: Library of Congress.

Card 4/4



MASEVICH, A.G.

Journey of Soviet astronomers to the United States. Astron. tsir.  
no. 184:24-26 S '57. (MIRA 11:4)  
(Russia--Relations (General) With United States)  
(United States--Relations (General) With Russia)

GINDIN, Ye.Z.; LNYKIN, G.A.; LOZINSKIY, A.M.; MASHKIN, A.G.; AL'PERT, Ya.L.;  
 CHUDSENKO, B.F.; SHAPIRO, B.S.; GALKIN, A.M.; GORLOV, O.G.; KOTOVA,  
 A.P.; KOSOV, I.I.; PETROV, A.V.; SEROV, A.D.; CHERNOV, V.N.;  
 YAKOVLEV, V.I.; MIKHAYLOV, A.A., otvetstvennyy red.; BEN'KOVA, N.P.,  
 doktor fiz.-mat. nauk, otvetstvennyy red.; SILKIN, B.I., red.;  
 PODOL'SKIY, A.D., red.; PRUSAKOVA, T.A., tekhn. red.

[Preliminary results of the scientific research on the first  
 Soviet artificial earth satellites and rockets; collection of  
 articles in the 11th section of the IGY program (rockets and  
 satellites)] Predvaritel'nye itogi nauchnykh issledovaniy s  
 pomoshch'yu pervykh sovetskikh iskusstvennykh sputnikov zemli  
 i raket; sbornik statei (XI razdel programmy MGO - rakety i  
 sputniki). Moskva, Izd-vo Akad. nauk SSSR, No.1. 1958. 148 p.  
 (MIRA 11:10)

1. Russia (1923- U.S.S.R.) Mezhdunarodnyy komitet po  
 provedeniyu Mezhdunarodnogo geofizicheskogo goda. 2. Chlen-kor-  
 respondent AN SSSR (for Mikhaylov).  
 (Atmosphere, Upper-Rocket observations)  
 (Artificial satellites)

MASEVICH, A.G.

X GENERAL ASSEMBLY OF THE INTERNATIONAL ASTRONOMICAL UNION

Moscow 12-26 Aug 58

Discussion on the Hertzsprung-Russell Diagram

EVOLUTION OF STARS DECREASING IN MASS

A.G. Mashevich

By a "steady" evolution of a star is understood such a change of its main parameters in the flow of time, which occurs in the range of an assumed equilibrium model and is caused by regular interior processes. So, for instance, the gradual conversion of hydrogen into helium in the convective core of a star, whose substance is not mixed fully, leads to a continuous change of its luminosity, radius,  $T_e$  and  $\rho$ , convective core dimensions. However, all these changes occur all the time in the limits of the model's state of equilibrium, and the structure of the star (radiative envelope and convective core) remains unchanged until a certain limit is reached in  $M$  ratio

$$\frac{M}{M_{\text{env}}} = \gamma_{\text{crit}} \quad /1,2,3/$$

After the value  $\gamma_{\text{crit}}$ , corresponding to the almost full exhaustion of hydrogen I in the convective core, is reached, further "steady" evolution becomes impossible, since for the further decrease of  $M$  in the range of the assumed model there are no

PHASE I BOOK EXPLOITATION

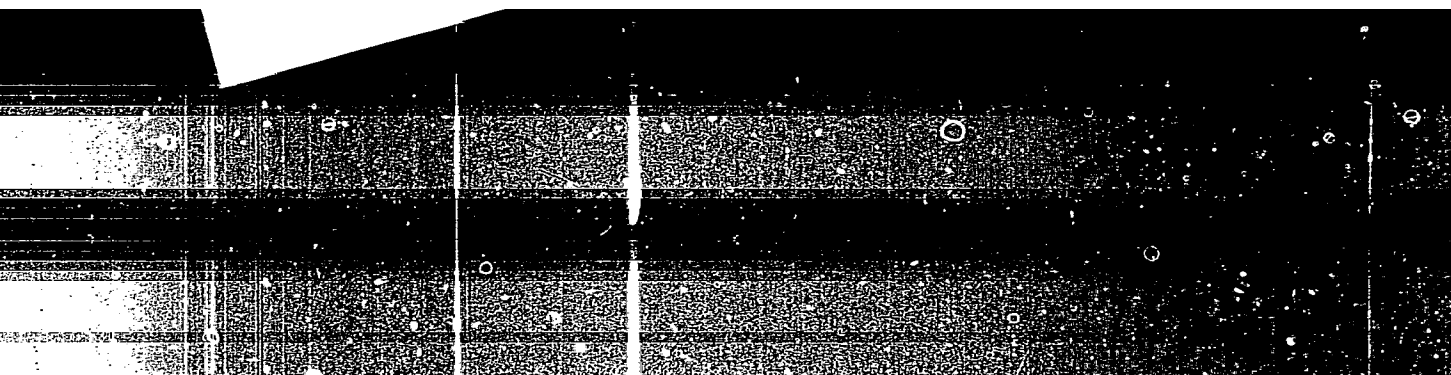
vesoyuznoye astronomo-geodezicheskoye obshchestvo  
(Astronomical Calendar; Yearbook. Variable Part; 1959)  
Fizmatgiz, 1958. 370 p. 8,500 copies printed. Moscow,  
Ed.: I.Ye. Rakhlin; Tech. Ed.: S.N. Akhlamov; Editorial Board:  
P.I. Bakulin (Resp. ed.), S.G. Kulagin, A.G. Masevich, and  
P.P. Parenago.

PURPOSE: This astronomical calendar is intended for specialists in  
astronomy, astrophysics, and geophysics.

COVERAGE: The book is divided into two parts. The first, based on  
data taken from the USSR Astronomical Yearbook for 1959, consists  
of ephemerides and accompanying text, compiled and written by the  
following specialists: S.G. Kulagin and L.D. Kovbasyuk of the  
State Astronomical and Geodetical Society) - notes on  
the ephemerides of the Sun and Moon; M.M. Dogayev  
of the Moscow Branch of the All-Union Astronomical and  
Geodetical Society) - text and maps of the visible trajectories of  
comets and maps of eclipses, the physical coordinates

"APPROVED FOR RELEASE: 07/12/2001

CIA-RDP86-00513R001032710013-3



APPROVED FOR RELEASE: 07/12/2001

CIA-RDP86-00513R001032710013-3"

MASEVICH, A.G.

3(1)

PHASE I BOOK EXPLOITATION

SOV/1840

Vsesoyuznoye astronomo-geodezicheskoye obshchestvo

Astronomicheskiy kalendar; yezhegodnik. Peremennaya chast'; 1959  
(Astronomical Calendar; Yearbook. Variable Part; 1959) Moscow,  
Fizmatgiz, 1958. 370 p. 8,500 copies printed.

Ed.: I.Ye. Rakhlin; Tech. Ed.: S.N. Akhlamov; Editorial Board:  
P.I. Bakulin (Resp. ed.), S.G. Kulagin, A.G. Masevich, and  
P.P. Parenago.

PURPOSE: This astronomical calendar is intended for specialists in  
astronomy, astrophysics, and geophysics.

COVERAGE: The book is divided into two parts. The first, based on  
data taken from the USSR Astronomical Yearbook for 1959, consists  
of ephemerides and accompanying text, compiled and written by the  
following specialists: S.G. Kulagin and L.D. Kovbasyuk of the  
GAGO (State Astronomical and Geodetical Society) - notes on  
ephemerides, the ephemerides of the Sun and Moon; M.M. Dogayev  
of the MOVAGO (Moscow Branch of the All-Union Astronomical and  
Geodetic Society) - text and maps of the visible trajectories of  
the planets, text and maps of eclipses, the physical coordinates

Card 1/10

Astronomical Calendar; Yearbook. Variable Part; 1959 SOV/1840

of the Sun, Moon, Mars, and Jupiter, the satellites of Jupiter and Saturn; N.D. Rozenblyum (MOVAGO) - ephemerides and heliocentric longitudes of planets; I.F. Yegorchenko, A.A. Kaverin, T.G. Konstantinova, V.A. Kuklina, G.V. Kuklin, Z.G. Sazonova, L.I. Chernykh, and N.S. Chernykh - data on 144 points in the USSR for the full solar eclipse of October 2, 1959; Ye.G. Demidovich (GAGO) - occultation of the stars and planets by the Moon, observation of the Polar Star, computation of stellar coordinates; V.A. Bronshteyn (MOVAGO) - comets; N.S. Yakhontova - the lesser planets; and, N.B. Perova (MOVAGO) - variable stars. The second part, the Supplement, contains a review of the achievements in astronomy for the years 1956 and 1957, written by V.A. Bronshteyn, O.D. Dokuchayeva, L.A. Katasev, M.A. Klyakotko, P.P. Parenago, and I.S. Shcherbina-Samoylova under the editorship of A.G. Masovich, articles on artificial satellites, the danger in astronautics from meteors, the nature of galaxies, articles on scientific meetings held in the Soviet Union and abroad, and articles on the anniversary of events in astronomy. The book is profusely illustrated with tables, maps, photographs, and diagrams. The Supplement includes some 125 Soviet references grouped according to subject matter and type of publication.

Card 2/10

Astronomical Calendar; Yearbook. Variable Part; 1959 SOV/1840

TABLE OF CONTENTS:

From the Editors 5

PART I. EPHEMERIDES

Explanations to the Ephemerides 7

Ephemerides of the Sun and Moon 16

Planets 40

Eclipses 64

Occultation of Stars and Planets by the Moon 79

Physical Coordinates of the Sun, Moon, Mars, and Jupiter 90

Jupiter's Satellites 98

Card 3/10



Astronomical Calendar; Yearbook. Variable Part; 1959	SOV/1840
Saturn's Satellites	
Comets	110
Lesser Planets	111
Variable Stars	111
Notes on Observations of the Polar Star	113
Notes on the Computation of Stellar Coordinates	121
	130

PART II. SUPPLEMENTS

Advances in Astronomy in the Years 1956 and 1957	134
This article discusses the observatory studies made on solar activity, the structure and temperature of the chromosphere, the exterior of the solar corona, studies conducted at the Crimean Astrophysical Observatory, large-scale and turbulent motions in the Sun's photosphere, studies of the Sun's general and localized magnetic fields, the stars	

Card 4/10

Astronomical Calendar; Yearbook. Variable Part; 1959

SOV/1840

including the variable ones, the spiral structure of the Galaxy, the Sun, the planets, comets, the Moon's atmosphere, the nature of Venus and Mars, and the meteors.

Artificial Satellites of the Earth and the Danger in Astronautics  
From Meteors (V.V. Fedynskiy)

197

The author reports mainly on studies of cosmic rays, the Sun's corpuscular radiation, micrometeorites (recorded by means of ammonium-phosphate piezoelectric counters) and the annual distribution of micrometeorites and their tentative quantities.

The Mrkos Comet (1957 d) (F.Yu. Zigel')

208

This article discusses the Mrkos Comet which was discovered on August 3, 1958. The comet's parabolic orbital elements are computed and the comet photographed. Observed by several Soviet astronomers its study provided much new material.

Card 5/10

- Astronomical Calendar; Yearbook. Variable Part; 1959 SOV/1840
- Noctilucent Clouds in 1957 (N.I. Grishin) 214  
Stereotriangulation methods for determining the height of clouds are described.
- Interaction and Nature of Galaxies (B.A. Vorontsov-Vel'yaminov) 231  
This article treats galactic bodies, tails, the units bridging them, and also double and multiple galaxies.
- Soviet Astronomers in the United States of America (A.G. Masevich) 243  
This article describes the June-July 1957 visit of a Soviet delegation of astronomers, headed by V.A. Ambartsumyan, to the United States.
- The Eighth International Astronautical Congress (A.G. Masevich) 263  
This article describes the Astronautical Congress held October 12, 1957 in Barcelona.

Card 6/10

Astronomical Calendar; Yearbook. Variable Part; 1959 SOV/1840

Joint Visiting Session of the Astronomical Council of the AN  
SSSR and the Academy of Sciences of the Azerbaydzhan SSR  
(M.A. Klyakotko)

271

This article treats the meeting at which M.M. Aliyev,  
A.A. Mikhaylov, A.A. Yakovkin, S.K. Vsekhsvyatskiy,  
V.V. Sharonov, V.P. Shcheglov, Z.I. Khalilov, V.A. Krat,  
and G.F. Sultanov participated.

The 350th Anniversary of the Formulation of Keppler's First  
Two Laws (Yu.A. Ryabov)

275

This article is a historical account and discussion of  
Keppler's first two Laws.

The 85th Anniversary of the Tashkent Astronomical Observatory  
(V.P. Sheglov)

286

The article provides a detailed historical account and  
description of the Tashkent Astronomical Observatory of  
the Academy of Sciences of the Uzbek SSR, the oldest scien-  
tific research institution in Central Asia. The Observatory

Card 7/10

Astronomical Calendar; Yearbook. Variable Part; 1959

SOV/1840

maintains its own meteorological station, a Time Station which provides 17 time signals in 24 hours, a Solar Laboratory which conducts systematic studies of the Sun's chromospheric flares on the basis of spectroscopic and photometric observations (Yu.M. Slonim, Chief, and K.F. Kuleshova, Z.B. Korobova, and B.N. Tirnshteyn, staff members), and a network of meteorological and other research stations. Of particular interest is the Kitaba International Latitude Station imeni Ulugbek situated 3 km. from the town of Kitaba in the Kashka-Dar'inskaya oblast'. Administered by the Observatory since 1941, the Station has conducted regular observations since 1930. Its staff members include A.M. Kalmykov, Director, D.I. Kravtsev, scientist, and P.V. Shcheglov and V.S. Obratsov, laboratory assistants. A zenith-telescope APM-2 was installed there in June 1958. In 1932 the Observatory came under the jurisdiction of the Committee on Science of the Central Executive Committee of the Uzbek SSR, since which time it has engaged in a program of research in exact time determination, solar activity, and meridian and photographic astronomy. It had been conducting regular observations of sun spots and solar protuberances since 1932. The Observatory's staff includes M.F. Bykov, who completed the work begun in 1945 of determining the direct ascension of weak stars by the absolute

Card 8/10

Astronomical Calendar; Yearbook. Variable Part; 1959

SOV/1840

method; Kh.R. Shakirova, B.V. Yasevich, and A. Kadyrov, who made thorough studies with two passage instruments of personal and instrument errors; V.P. Shcheglov, V.T. Beda, B.Zh. Bal'zhinova, B.V. Yasevich, N.A. Omelina, L.N. Koshkina, M.G. L'vova, and G.I. Kazakov, who, in cooperation with IGY program, engaged in daily determinations of time corrections on two passage instruments and in the reception of a large number of rhythmic signals, V.A. Mal'tsev and N.N. Sytinskaya - observation of meteors; A.A. Latypov, I.M. Ishchenko, and G. Kim - regular photographic observations of the Earth's artificial satellites; F.G. Ustimenko, Chief Mechanical Engineer, and Ye.P. Kolesnikova, Head Librarian. Some of the newer equipment possessed by the Observatory include: a passage instrument APM-10, new printing chromographs, radio reception and measurement apparatus, two sets of quartz clocks obtained in 1958, a normal astrograph, a meridian circle, a zenith-telescope APM-2 set up in 1957, a solar protuberance spectroscop (obtained 1932), a standard spectroheliometer (obtained 1935), a

Card 9/10

Astronomical Calendar; Yearbook. Variable Part; 1959 SOV/1840

chromosphere-photosphere telescope, a celostat with a clock mechanism for spectrohelioscope, and a microphotometer MF-4. The Tashkent Astronomical Observatory (TAO) published its own Trudy, a Byulleten', and Circulars.

The 70th Anniversary of the Gor'kiy Division of the All-Union Astronomical-Geodetical Society (S.G. Kulagin) 315

Anniversary of Soviet and World Astronomy in 1959 (Yu.G. Perel') 325  
The article treats briefly the Committee on Solar Research of the Academy of Sciences, USSR.

The Tenth International Astronomical Meeting in Moscow (D.Ya. Martynov) 350

Bibliography (Yu.G. Perel') 362

AVAILABLE: Library of Congress

Card 10/10

MM/ad  
6-17-59

80794

SOV/169-59-6-6375

3.2300

Translation from: Referativnyy zhurnal, Geofizika, 1959, Nr 6, pp 140 - 141 (USSR)

AUTHORS: Gindin, Ye.Z., Leykin, G.A., Lozinskiy, A.M., Masevich, A.G.

TITLE: The Optical Observations of Artificial Earth Satellites ✓

PERIODICAL: V sb.: Predvarit. itogi nauchn. issled. s pomoshch'yu pervykh sov. iskustv. sputnikov Zemli i raket, Moscow, AS USSR, 1958, pp 5 - 39 (Engl. Res.)

ABSTRACT: The Astronomicheskii sovet Akademii nauk SSSR (Council of Astronomy of the USSR Academy of Sciences) was put in charge of organizing the optical observations of artificial earth satellites. Sixty-six visual stations and twenty-four photographic stations were established for observing the satellites. The visual observation stations began their activity at the time when the first Soviet satellite was launched, while photographic observations have been performed systematically since the beginning of 1958. The visual observation methods were determined by the task: they must establish the position of a satellite or

Card 1/6



80794  
SOV/169-59-6-6375

# The Optical Observations of Artificial Earth Satellites

the celestial sphere with an accuracy of  $0.5$  to  $1^\circ$  and the time within  $0.5$  to  $1$  sec, and must report the observation results to the computer center within the shortest time. Two "optical barriers", each consisting of about 30 telescopes, were established to facilitate the observation of satellites having a low brightness and moving on the sky with a velocity of  $1^\circ$  per  $1$  sec, if the orbit is known only approximately. The barriers are located on the meridian and along a vertical circle perpendicular to the visible orbit of the satellite. The sight lines of the telescopes are adjusted in such a way that each section of the optical barrier is covered twice. For determining the time of passage of a satellite with an accuracy exceeding  $1$  sec, the time signals and the signals given by the observer at the time when the satellite passed, are recorded on tape. After the termination of the observations, the tape recording is reproduced at a low speed and the precise moment of passage is determined by a chronoscope. The coordinates of the satellite are determined by the sidereal maps of A.A. Mikhaylov's atlas or of A. Bechvarzh's atlas. When observing satellites of low brightness ( $15 - 8$  stellar magnitude) the AT-1 telescope is used, which is a small wide-angle telescope having a  $50$  mm objective lens and six-power magnification. The field of view is  $11^\circ$

Card 2/6

80794

SOV/169-59-6-6375

# The Optical Observations of Artificial Earth Satellites

The stations observing the satellites are provided with signals of the correct time by feeding to them second tone signals. On the basis of observation data, the computer center informs the stations on the coming passage of a satellite. The station receives a coded telegram containing information on the time and altitude of a satellite's passage in the meridional plane and in the plane in which the nearest point of the orbit is located. Observations of artificial satellites are also performed on the territories of the Chinese People's Republic (KNR), the German Democratic Republic (GDR), Czechoslovakia, Poland, Hungary, Rumania, and Bulgaria, where 45 stations are in operation. Further, observatories in England, Scotland, Ireland, the US and other countries were included in the visual and photographic observation system of the Soviet satellites. At some stations, besides the visual observations, the positions of the carrier rocket and the second Soviet satellite are determined photographically by "Zorkiy" cameras with "Yupiter 8" lenses.<sup>o</sup> At the time of the satellites passage across the field of view of the camera, the shutter is opened for a brief time interval (2 - 5 sec). The begin and the end of the exposure are marked by a chronograph. It is possible to determine by photo-

Card 3/6

80794

SOV/169-59-6-6375

# The Optical Observations of Artificial Earth Satellites

graphic observations the position of a satellite with an accuracy of  $3' - 5'$  of arc. The Council of Astronomy discussed the problem of using light flashes of short duration on the object for a precise determination of a satellite's position. The position of a satellite may be determined with an accuracy of 2 - 3 sec of arc when using cameras with a long focal length ( $F \sim 1$  m) for photographing the satellite. Using the data of these observations for triangulation on the earth's surface, the distance between different points (especially between continents) and also the shape of the geoid may be determined with an accuracy of 10 m. However, the photography of satellites is made difficult by the following circumstances: 1) the observations are possible only at dusk; 2) cameras with a very great light power are required; 3) the setting of precise time marks is complicated. These difficulties can be overcome if the satellite is equipped with a light source producing brief flashes by which it may be photographed at night. It is expedient to provide series of flashes and not a continuous feed, taking into consideration that at least two or three flashes must arrive in the field of view of the instrument. In this way it is possible to determine not only the position but also the angular velocity of a satellite. Obviously,

Card 4/6

80794

SOV/169-59-6-6375

# The Optical Observations of Artificial Earth Satellites

a pulse gas discharge lamp should be used as a light source, whose light output reaches 60 lm/w. The brightness of a satellite depends on the following reasons: 1) changes in the satellite's phases, i.e. in the configuration sun - satellite - observer; 2) changes in the distance to the observer; 3) light absorption in the section of its path from the satellite to the observer; 4) rotation and tumbling of a satellite, 5) changes in the state of the satellite's surface. The determination of the period of rotation (tumbling) of the satellite's body and changes of this period in time are of the greatest interest. Another important problem is the investigation of the dependence of the brightness and color of a satellite on the state of the earth's atmosphere. Finally, the third problem is the change of the state of the satellite's surface under the influence of the atmosphere and extraterrestrial agents. For solving the aforementioned problems a precise quantitative determination of brightness changes of a satellite and observations over a possibly great section of its trajectory are necessary. Presently, two methods are used for measuring a satellite's brightness. The first method consists in a

44

Card 5/6

80794

SOV/169-59-6-6375

The Optical Observations of Artificial Earth Satellites

comparison of the brightness of the satellite's trail with the brightness of the trails of neighboring stars on a photography obtained by a stationary camera. The second method consists in a visual comparison of the satellite's brightness with the brightness of stars located along its path. Both methods are used at Soviet observation stations. *UH*

L.V. Terent'yeva

Card 6/6

LOZINSKIY, A.M.; MASEVICH, A.G.

Optical observations of artificial earth satellites in the U.S.S.R.  
Meshdunar. geofiz. god no.5:23-28 '58. (MIRA 11:10)  
(Artificial satellites)

**AUTHOR:** Vasevich, A. A., Doctor of Science, Professor, Institute of  
**TITLE:** An Important Contribution to the Development of Biological  
 Science (Kryazev's vivid & revealing analysis of the work of  
 The 15th International Congress on the History and Philosophy of  
 International Space Astronomy)  
**PERIODICAL:** Priroda, 1968, No 10, pp 15-21, 15 figs.  
**ABSTRACT:** The Tenth International Congress on the History and Philosophy of  
 Science held in Moscow from 21 August 1968 was attended by over 1000 people  
 and 300 Soviet participants and guests from 45 countries.  
 The Congress was opened by its Secretary A. A. Vasevich. The  
 participants in the Congress were welcomed by the Deputy  
 President of the USSR Council of Ministers, A. A. Gromyko,  
 who pointed out the great achievements of Soviet science  
 and stated that the Soviet Government is fully aware of the  
 importance of international scientific cooperation. The  
 Congress was supported by the President of the USSR Academy of Sciences,  
 of the Congress, Academician N. S. Keldysh, opened the Tenth  
 President of the USSR Academy of Sciences, Academician  
 spoke. The main themes concerned the evolution of the  
 the origin of the chemical elements, the origin of life, the

An Important Contribution to the Development of Astronomical  
Xth International Congress of Astronomers

of the earth and the stars, frequently standard astronomical observations by means of balloons, rockets and artificial satellites, the sun and its influence upon the earth, the origin of the solar system, comets and meteors, a new theory on the origin of the universe, which was put forward by several American scientists, was mentioned by the Soviet physicist Andrei Sakharov. He stated that the hypothesis presupposed the existence of a large number of "dead" stars in the universe, which are formed by the fusion of lighter nuclei can be formed also on the surface of stars, in the so-called active zones with magnetic fields, capable to accelerate the smallest particles. This assumption has been supported by observations of energetic particles obtained in the Kryazevskaya astronomical observatory. It is possible that the same processes might be occurring in the explosions occurring on the surface of stars. The presence of the active zones on the surface of stars

Card 2.4



An Important Contribution to the Development of Astronomical Science  
Xth International Congress of Astronomers

atmospheres of certain stars would also be explained by  
Kel'dovich's suggestion. Soviet astronomer L. V. Keldysh  
has analyzed over 20,000 observations of the movement of the  
geographical pole of the earth and collected the complete  
part of the earth has the properties of a liquid body. Soviet  
scientists L. N. Vernov and A. I. Zhuravskiy presented the results  
of the study of cosmic rays and proton by means of the  
second and third satellites. Interesting observations of the  
magnetic field of the sun were carried out by L. V. Keldysh  
and Vernov, by means of the special spectrograph of the Soviet  
Astrophysical Observatory. In addition to the results of the  
Soviet and Czech astronomers collected at the time of the  
connection between solar activity and geophysical processes  
on earth. Thus the highest level of the geomagnetic  
activity coincides with the secular minimum of solar activity.  
The results of the international study of the variable stars  
have been collected, arranged and published in the book.  
During the Congress, the scientific edition of the book was  
containing the description of 10,000 variable stars.

Card 3/4

An Important Contribution to the Development of Astronomical Science  
Xth International Congress of Astronomers.

Five numbers of the Congress periodical in "English, Russian,  
languages" were published at this time. The delegates  
visited scientific institutions in Moscow and in several  
astronomical observatories in the Crimea, Armenia, Georgia  
and Uzbekistan.

There are 1 photo and 1 graph.

Card 4/4

*11/11/1958*  
 AUTHOR: Masevich, A. G., Deput. Chairman of the Astronomical Council AS USSR 10-1-3/10

TITLE: A New Efficacious Instrument for the  
**Knowledge of the Universe** (Novoye effektivnoye orudiye  
 poznaniya vseleeniya).

PERIODICAL: Vestnik AN SSSR, 1958, Vol. 26, Nr 1, pp. 8-12 (USSR)

ABSTRACT: The first earth satellite was of spherical shape and had a diameter of 58 cm and a weight of 83,6 kg. It was launched on October 4, 1957 and disappeared on January 4, 1958. The greatest height of its orbit was 900 km. The second earth satellite forms the last rocket step and weights more than 1/2 ton. The greatest height of its orbit is 1700 km. The orbits of the earth satellite have approximately the form of an ellipse, one of the foci of which is in the center of the earth. The plane of both orbits is inclined towards the equatorial plane at an angle of 65°. Because of the rotation of the earth, the earth satellite in each successive revolution passes a domain which is located approximately by 1000 to 2000 km more West than the preceding one, in dependence of the degree of latitude.

Card 1/3

A New Efficacious Instrument for the Knowledge of  
the Universe

10-1-2/39

The period of a rotation of the first earth satellite at first amounted to 96,2 minutes and that of the second to 103,7 minutes. The earth satellites are visible only at dusk without the help of a telescope etc. For the purpose of optical observation of the artificial earth satellites, the Astronomical Council AN USSR organized 66 special stations at the Physical-Mathematical Faculties of universities and pedagogic institutes. A special small-sized telescope AT-1 was developed and produced, of which each station received 20-25; furthermore, star charts, stop watches, and a special device for recording observed times and the reception of exact time signals were provided. The organization for observation is described in detail, in which connection also the photographic method is mentioned. On the basis of deviations and modifications of the satellite orbit conclusions may be drawn concerning the modification of the force of gravity at various points of the earth. In this manner the distribution of the mass of the earth can be ascertained, which is not possible in a thorough manner by any other method. The artificial earth satellites are

Card 2/3

A New Efficacious Instrument for the **Knowledge of**  
the Universe

20-1-3/59

flying laboratories and make direct observations of the ionosphere of the upper atmospheric strata, of shortwave-solar radiation, possible, which is impossible to be carried out from the earth. In this way also the primary cosmic rays can be investigated. By taking with it a living being (dog) the second artificial earth satellite made it possible to carry out medical and biological investigations, the results of which will contribute greatly towards making the flight of human beings into cosmic space without danger possible. In this way the most important problems of astro- and geophysics can be investigated and studied, which is impossible on earth.

ASSOCIATION: Astronomical Council AN USSR (Astronomicheskii sovet Akademii nauk SSSR).

AVAILABLE: Library of Congress

Card 3/3

1. Satellites-Nomenclature
2. Satellites-Motion
3. Satellites-Observation

3(1)

AUTHOR: ~~Masovich, A.G.~~

SOV/33-35-2-16/21

TITLE: On Some Peculiar Stars Below the Main Sequence of the H-R Diagram (O nekotorykh pekuliarnykh zvezdakh, nakhodyashchikhsya pod glavnoy posledovatel'nost'yu diagrammy Ressella)

PERIODICAL: Astronomicheskii zhurnal, 1958, Vol 35, Nr 2, pp 292-294 (USSR)

ABSTRACT: According to the proposal of P.P.Parenago the author considers four stars ( $\lambda$  Boo, 29 Cyg,  $\gamma$  Aqr, 2 And) with peculiar spectra which lie below the main sequence of the H-R diagram. There seems to be no reason for taking these stars as representatives of population II. Their present position on the H-R diagram can be explained if they are considered as initially main sequence stars ( $1 \odot < M < 2 \odot$ ,  $0.8 < X < 0.9$ ,  $0.01 < Z < 0.1$ ) which evolved off the main sequence as a result of an evolution with constant mass and complete mixing till a stage corresponding to X-0.5 (fig 1). The time of such an evolution is about  $3 \cdot 10^9 < t < 8 \cdot 10^9$  years.

There are 2 tables, 1 figure, and 5 references, 3 of which are Soviet, and 2 American.

ASSOCIATION: Gosudarstvennyi astronomicheskii institut imeni P.K.Shternberga (State Astronomical Institute imeni P.K.Shternberg)

SUBMITTED: January 26, 1958

Card 1/1

3(1)

AUTHOR: Masevich, A.

SCV/33-35-6-17/18

TITLE: The 10-th International Astronomical Congress in Moscow

PERIODICAL: Astronomicheskiy zhurnal, 1958, Vol 35, Nr 6, pp 941-956 (USSR)

ABSTRACT: The paper contains a report on the Astronomical Congress which took place from August 12 - 21, 1958 in Moscow. There were 1209 participators from 35 countries. The Soviet delegation consisted of 411 scientists, from the USA there participated 225, from France 103, from Germany 73, and from England 63 scientists etc.

The opening session was addressed by A.I. Kosygin, Vice-President of the Council of Ministers of the SU by A. A. Mikhaylov, President of the Astronomical Council of the Academy of Sciences of the USSR. Furthermore, A.V. Topchiyev, Vice-President of the Academy of Sciences, and V.A. Ambartsumyan, President of the Organization Committee of the Congress, also spoke.

A journal of the congress with the title "Kosmos" was published.

The following films were shown : 1. "The sun in action", USA,

Card 1/2

The 10-th International Astronomical Congress  
in Moscow

NOV/33-35-6-17/18

subtitles in Russian by S. Gaposhkin. 2. "Soviet observatories" by B.A. Vorontsov - Vel'yaminov etc. The following Soviet scientists gave lectures : P.P. Parenago, I.M. Kopylov, P.N. Kholopov, K.A. Barkhatova, V.A. Ambartsumyan, A.G. Masevich, V.A. Krat, E.R. Mustel', V.G. Fesenkov, G.M. Idlis, Ya.I. Fedorov, A.A. Nemiro, N.N. Pavlov, D.A. Frank - Kamenetskiy, P.E. Nemirovskiy, B.A. Tverskiy, R.Z. Sagdeyev, A.B. Severnyy, S.N. Vernov, A.E. Chudakov, T.N. Nazarova, V.I. Krassovskiy, Ya.L. Al'pert, L.A. Zhekulin, A.N. Kazantsev, V.P. Tsesevich, Ye. Pavlovskaya, Yu. Pskovskiy, S.A. Zhevakin, G.A. Idlis, O.M. Mel'nikov, A.P. Vinogradov, A.A. Yavnel', L.G. Kvash, Ye.L. Ruskol, A.I. Lebedinskiy, B.Yu. Levin, V.S. Safronov, S.K. Vsekhsvyatskiy, K.A. Shteyns, V.I. Cherednichenko, I.S. Astapovich, A.L. Zel'manov, A. Masevich, P.V. Shcheglov, V.B. Nikonov, M.S. Zverev, Ye.P. Fedorov, Ye.K. Kharadze, G.D. Mamedbeyli, A.N. Deych, and M.S. Eygenzon.

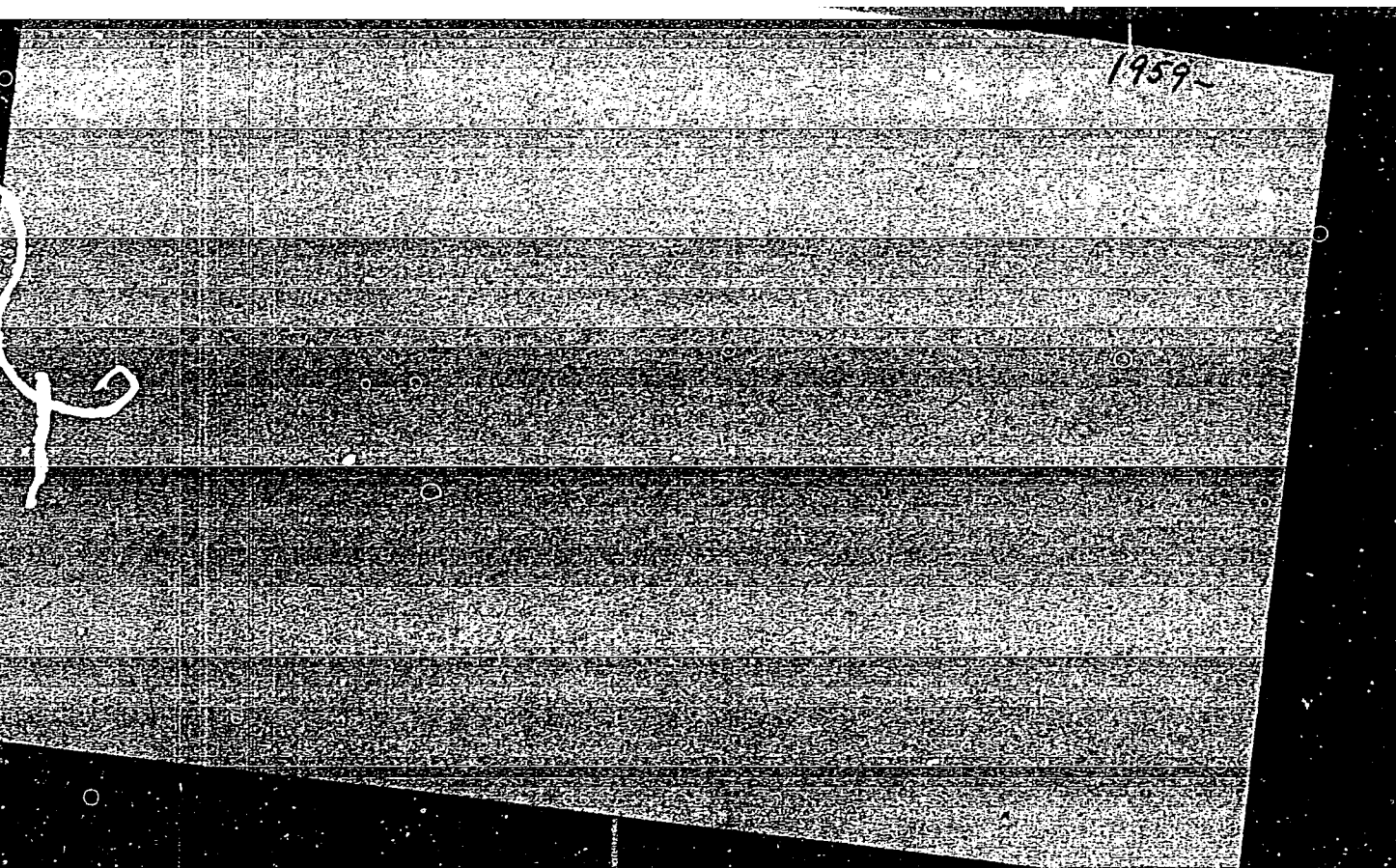
SUBMITTED: October 3, 1958

Card 2/2



"APPROVED FOR RELEASE: 07/12/2001

CIA-RDP86-00513R001032710013-3



APPROVED FOR RELEASE: 07/12/2001

CIA-RDP86-00513R001032710013-3"



MASEVICH, Alla G.

"Model of the Sun."

report to be submitted for the 9th Intl. Symposium, Belgian Inst. of Astrophysics,  
Liege, Belgium, 6-8 July 1959.

66609

~~3(1), 29(0)~~ 3.2100

AUTHOR:

Masevich, A. G., Doctor of Physical and  
Mathematical Sciences

SOV/30-59-5-21/43

TITLE:

Astronomical Observations of Artificial Earth Satellites  
(Astronomicheskiye nablyudeniya iskusstvennykh sputnikov Zemli).

PERIODICAL:

Vestnik Akademii nauk SSSR, 1959, Nr 5, pp 85-94 (USSR)

ABSTRACT:

Observations are carried out in the USSR by means of the telescope AT-1 the apparatus of which was recently improved by B. Ye. Tumanyan (Yerevan). Several stations use theodolites fitted with telescopes AT-1 (Khabarovsk, Kyzl-Orda, Tartu). On this basis J. E. Elnasto (Tartu) constructed an experimental automatic recording theodolite. D. Ye. Schchegolev investigated at the station of the Pulkovo observatory the methods of observing artificial earth satellites. Many stations photograph the earth satellites with the narrow film camera of the "Kiyev" and "Zorkiy" type. In the South-Sakhalin station K. N. Kan used for this purpose the telescope AT-1 as a sighting telescope. The station of the Vologda Pedagogical Institute (A. P. Poletayev) achieved considerable success in this field. In the USSR the artificial earth satellites are taken with wide-angle aerial photography cameras NAFA-3s/25

Card 1/3

66609

**Astronomical Observations of Artificial Earth Satellites** SOV/30-59-5-21/43

with "Uran 9" lenses. A great part of the stations determine the position of the earth satellites by means of the method of A. N. Deych and A. A. Kiselev (Pulkovo). In order to secure still more accurate coordination of the earth satellites large astronomical telescopes are provided with a device which permits an exact recording of the time of exposure; this device was for the first time constructed in the Astrofizicheskii institut Akademii nauk Kazakhskoy SSR (Astrophysics Institute of the Academy of Sciences of the Kazakh SSR). In this connection the author of the present paper refers to the paper by D. A. Rozhkovskiy in this periodical (foot note 2). The Astronomical Observatory imeni Engel'gardt near Kazan' constructed a special plate holder for telescopes which permits the passage through certain points of the photo plate to be recorded with respect to time. Ye. Ya. Bugoslavskaya works out a similar method of photographing earth satellites in the Astronomical Institute imeni Shternberg in Moscow. L. A. Panaytov of the Astronomical Main Observatory (Pulkovo) built a cinematographic camera. M. K. Abele, scientific collaborator of the Riga station suggested an interesting method. The Odessa Observatory evaluated results of observations made at different

Card 2/3

66609

**Astronomical Observations of Artificial Earth Satellites** SOV/30-59-5-21/43

points of the surface of the earth. The author of the present paper mentions further the papers by V. P. Tsesevich (Odessa) and M. S. Zverev as well as M. I. Yesipova (Pulkovo). The results of observations of the motion of artificial earth satellites show how much the scientists all over the world endeavor to collaborate objectively. An insert shows photographs of the three Soviet artificial earth satellites in the years 1957 and 1958. All these observations are important because the orbits of the earth satellites are still within the earth's atmosphere and because their deviations from the orbits computed according to the laws of celestial mechanics permit conclusions as to the composition of atmosphere. There are 13 figures and 2 references, 1 of which is Soviet.

4

Card 3/3

3(1)  
 AUTHOR: Masevich, A. G., Doctor of Physical and SOV. 29-12-1959, 3  
~~Mathematical Sciences~~, Deputy Chairman of the Astronomic  
 Council AS USSR

TITLE: New Cosmic Experiments Are Planned

PERIODICAL: Tekhnika molodezhi, 1959, Nr 11, p 9 (USSR)

ABSTRACT: This is a short consideration of the purpose and use of  
 advance into cosmos. The scientific research flights made in  
 recent years converted astronomy from a deliberative into an  
 experimental science. The possibility of observing celestial  
 bodies outside the terrestrial atmosphere is of utmost impor-  
 tance. Astronomers obtained many new data on the Moon, <sup>also</sup>  
 by taking photographs of the Far Side. Further possibilites ✓  
 will include the experimental investigation of Mars, Venus  
 and other planets in the solar system. It is hoped that the  
 desire of astronomers to establish an observatory on the Moon ✓  
 will be realized. Many new discoveries can be made with a  
 telescope in the cosmic space. The investigation of ultraviolet  
 rays absorbed by the terrestrial atmosphere might give valuable  
 hints as to processes taking place on various celestial bodies.

Card 1/1

ASSOCIATION: Astronomicheskii soviet AN SSSR (Astronomic Council, AS USSR)

20

3(1)  
AUTHOR: Masevich, A.G. SOV/33-36-3-29/29  
TITLE: 101<sup>st</sup> Congress of the American Astronomical Society  
PERIODICAL: Astronomicheskiy zhurnal, 1959, Vol 36, Nr 3, pp 557-560 (USSR)  
ABSTRACT: This is a report on the annual congress of the American  
Astronomical Society, December 29-30, 1958 in  
Florida. It contains a short index of the lectures and a photo  
of the participators of the congress.  
There is 1 figure.  
SUBMITTED: March 18, 1959

Card 1/1



MASEVICH, Alla Genrikhovna, doktor fiz.-mat.nauk

Secrets of the moon. Rabotnik 37 no.12:27 D '59.  
(MIRA 13:3)

1. Zamestitel' predsedatelya Astronomicheskogo soveta  
AN SSSR.  
(Moon--Surface)

DYBOVSKAYA, Irma Konstantinovna, dotsent, kand.filol.nauk; PROMTOVA, Irina Andreyevna; SUVOROVA, Vera Vasil'yevna; CHESKIS, Zoya Borisovna; DEYEV, G.N., red.; MASEVICH, A.G., doktor fiz.-matem.nauk, red.; PARIYSKIY, N.N., kand.fiz.-matem.nauk, red.; TANTSOVA, N.N., kand. tekhn.nauk, red.; TERENT'YEVA, L.V., red.; TYAGUNOVA, Z.I., red.; KRYUCHKOVA, V.N., tekhn.red.

[French-Russian geophysical dictionary] Frantsuzsko-russkii geofizicheskii slovar'. Pod red. G.N.Deeva i dr. Moskva, Glav.re-daktsiia inostr.nauchno-tekhn.slovarei Fizmatgiz, 1960. 374 p.  
(Geophysics--Dictionaries) (MIRA 13:9)  
(French language--Dictionaries--Russian language)  
(Russian language--Dictionaries--French language)

MAS VICH, H. 10.

PLATE I BOOK EXPLOITATION

50V/996

Mitshaylov, A. A., ed.

Stanislav y kosmos: Ispolnizatsiya (Space Stations: Collection of Articles) Moscow, Izd-vo AN SSSR, 1960. 144 p. 25,000 copies printed. (Series: Nauchno-populyarnaya knizhnitsa)

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. This book is intended both for the space specialist and the average reader interested in space problems.

Comments: The book contains 73 short articles by various Soviet authors on problems connected with space travel and the launching of artificial earth satellites and space rockets. Some possibilities of future developments are also discussed. The articles were published in the period of 1957-1960. No personalities are mentioned. There are no references.

~~... ..~~  
Mitshaylov, A. A., ed. Doctor of Physical and Mathematical Sciences. From the Earth to the Moon [September 15, 1959] 272

Soviet Artificial Earth Satellites [Pravda, October 9, 1957] 76

Mitshaylov, A. A., ed. Candidate of Physical and Mathematical Sciences. Automatic Laboratory in Space [November 14, 1957] 90

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Investigation of the Upper Atmosphere with the Help of the Artificial Earth Satellite [October 10, 1957] 93

Soviet Artificial Earth Satellites [Pravda, April 27, 1958] 96

Mitshaylov, A. A., ed. Candidate of Physical and Mathematical Sciences. On the Way to an Understanding of the Universe [December 9, 1957] 112

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Corresponding Member of the Academy of Sciences USSR and L. I. Kurmosova, Candidate of Physical and Mathematical Sciences. The Sun, Cosmic Radiation, and Sputniks [November 14, 1957] 115

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Investigation of Outer Space [December 11, 1957] 118

Third Soviet Artificial Earth Satellite [Pravda, May 10, 1958] 124

Discovered, Widening Knowledge About the Universe [Pravda, October 5, 1958] 153

Mitshaylov, A. A., ed. Candidate of Physical and Mathematical Sciences. In Outer Space - Our Third Sputnik [July 1958] 174

Mitshaylov, A. A., ed. Doctor of Physical and Mathematical Sciences. Let's Look into Outer Space [March 22, 1956, December 11, 1957] 183

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Sputnik on a Photo Plate [March 1958] 186

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Doctor of Physical and Mathematical Sciences. Secret of the Mysteries of the Universe [May 18, 1958] 190

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Candidate of Physical and Mathematical Sciences. Does the Aurora of Reflected Light from the Sputniks Change? [September 12, 1958] 191

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. High Altitude Laboratories [May 16, 1958] 192

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Doctor of Physical and Mathematical Sciences. Outer Space Laboratory [1958] 194

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Corresponding Member of the Academy of Sciences USSR. Ascent on Outer Space [1958] 204

Repr. Ed. I. A. A. Mitshaylov; Compiler V. V. Pedorov; Ed. of Publishing House: Ye. M. Klyus; Tech. Ed. I. D. Novichova. Candidate of Biological Sciences. Life on

KHOLLICHER, Val'ter [Hollitscher, Walter]; AKCHURIN, I.A. [translator];  
ARKHANGEL'SKIY, N.S. [translator]; MOCHALIN, D.N. [translator];  
OMEL'YANOVSKIY, M.E., akademik, red.; OPARIN, A.I., akademik, red.;  
MASEVICH, A.G., doktor fiziko-matem.nauk, red.; OVCHINNIKOV, M.F.,  
kand.filosof.nauk, red.; TYURYUKANOV, A.N., kand.biolog.nauk, red.;  
GAL'PERIN, P.Ya., dotsent, red.; URYSON, M.I., kand.biolog.nauk,  
red.; MAKAROV, A.A., red.isd-va; ZOTOVA, N.V., tekhn.red.

[Nature in the scientific picture of the world] Priroda v nauchnoi  
kartine mira. Obshchaya red. i vstupitel'naya stat'ya M.E.  
Omel'yanovskogo. Moskva, Izd-vo inostr.lit-ry, 1960. 469 p.  
(MIRA 14:3)

1. AN USSR (for Omel'yanovskiy).  
(Science--Philosophy)

GINDIN, Ye.Z.; LEYKIN, G.A.; LOZINSKIY, A.M.; LUR'YE, M.A.; MASEVICH,  
A.G.; SEVERNAYA, O.A.; SENTSOVA, Yu.Ye.; SLOVOKHOTOVA, N.P.;  
TOL'SKAYA, V.A.; TSITOVICH, V.V.

Brief report of the Astronomical Council of the Academy of  
Sciences of the U.S.S.R. on visual and photographic observations  
of artificial earth satellites in 1957-1959. Biul. sta. opt.  
nabl. isk. sput. Zem. no. 6:1-33 '60. (MIRA 14:2)  
(Artificial satellites--Tracking)

FEDOROV, Ye.; MASEVICH, A., doktor fiz.-mat.nauk

Steps toward outer space. Tekh.mol. 28 no.6:8-9 '60.  
(MIRA 13:7)

1. Chlen-korrespondent AN SSSR (for Fedorov). 2. Zamestitel'  
predsedatelya Astronomicheskogo soveta AN SSSR (for Masevich).  
(Space ships)



78006  
307/51-11-1-1/1

Step 1. The of the Sun

relative intensity of the two energy sources. No  
mixing occurs between the different parts. Numerical  
integration of the model is performed from the surface  
to the interior and every step is checked for possible  
convective instability. The absorption law is written

$$\kappa = \frac{1}{\rho} \left( \kappa_{\text{H}} + \kappa_{\text{He}} + \kappa_{\text{H}} \frac{30(1+X)}{1+Y} \right) \quad (1)$$

$$\epsilon = 3.9 \cdot 10^{25} (1+X)(1-X-Y) + 4.1 \cdot 10^{27} (1-X)(Y+Y) \quad (2)$$

where  $\rho$  is density;  $t$  is correction factor taken  
from Monod's tables;  $X$  and  $Y$  are relative amounts of  
hydrogen and helium. A homogeneous model could be  
obtained only with  $M = M_{\odot}$ ,  $R = R_{\odot}$ ,  $L = L_{\odot}$ .  
( $M_{\odot} = 1.99 \cdot 10^{33}$  g,  $R_{\odot} = 6.96 \cdot 10^8$  cm,  $L_{\odot} = 3.84 \cdot 10^{33}$  erg/sec).  
and  $X = 0.75$  is the relative amount of helium  
elementary. This model has small convective core  
more than  $M_{\odot} = 1.99 \cdot 10^{33}$  g,  $R_{\odot} = 6.96 \cdot 10^8$  cm,  $L_{\odot} = 3.84 \cdot 10^{33}$  erg/sec

card 7/4



Structure of the Sun

1960  
SOV/11-11-11/11

other parts radiative equilibrium remains stable. The central temperature and density are:  $T_c = 12.67 \times 10^6$  °K and  $\rho = 150 \text{ g/cm}^3$ . The upper boundary of the intermediate zone is at  $r = 0.75 R_\odot$ , and the energy source of the model is the proton-proton reaction; even inside the core the contribution of the carbon cycle is moderate. The authors also carried out calculations for the chemical composition:  $X = 0.74$ ,  $Y = 0.25$ , and  $Z = 0.0075$ , assumed by P. Naur, and they obtained results similar to his. A table gives a summary of 15 different models computed by various authors from 1947 to 1959 which differ in assumed laws of absorption and in chosen chemical composition. There are 2 figures; 3 tables; and 21 references; 5 Soviet, 1 Chinese, 2 German, 14 U.S. The most recent U.S. references are: M. Schwarzschild, R. Howard, R. Harm, Astrophys. J., 125, 255, 1957; O. Abell, Astrophys. J., 121, 430, 1955; I. Epstein, R. Motz, Astrophys. J., 117, 311, 1953; Ph. Morse, Astrophys. J., 92, 27, 1940; R. L. Sears, Astron. J., 55, 53, 1953; Astrophys. J.,

Card 3/4

Structure of the Sun

7-001  
S07/33-57-1-6/11

ASSOCIATION:

1.0, 489, 1999.  
Sternberg State Astronomical Institute and Computing  
Center of Moscow State University (Gos. astronomicheskiiy  
In-t imeni P. K. Shternberga, Vychislitel'nyy tsentr MSU)

SUBMITTED:

September 1, 1999

Card 4/4

KULAGIN, S.G.; KOVBASYUK, L.D.; DAGAYEV, M.M.; LAZAREVSKIY, V.S.;  
 DEMIDOVICH, Ye.G.; BRONSHTEIN, V.A.; YAKHONTOVA, N.S. (Leningrad);  
 KUROCHKIN, N.Ye.; DOKUCHAYEVA, O.D.; SHCHERBINA-SAMOYLOVA, I.S.;  
 MASEVICH, A.G.; LIPSKIY, Yu.N.; MARTYNOV, D.Ya.; ARSENT'YEV, V.V.;  
 MOROZ, V.I.; MASEVICH, A.G.; PEREL', Yu.G.; BAKULIN, P.I., otv.  
 red.; KULIKOV, G.S., red.; AKHLAMOV, S.N., tekhn. red.

[Astronomical calendar; yearbook. Variable part, 1962] Astronomicheskii kalendar'; ezhegodnik. Peremennaya chast', 1962. Red. kollegiya: P.I. Bakulin i dr. Moskva, Gos. izd-vo fiziko-matem. lit-ry, 1961. 259 p. (Vsesoyuznoe astronomo-geodezicheskoe obshchestvo, no. 65) (MIRA 14:12)

1. Gosudarstvennoye astronomo-geodezicheskoye obshchestvo (for Kalugin, Kovbasyuk, Lazarevskiy, Demidovich). 2. Moskovskoye ot-deleniye Vsesoyuznogo astronomo-geodezicheskogo obshchestva (for Dagayev, Bronshten, Kurochkin).  
 (Astronomy—Yearbooks)

MASEVICH, A. G. (Prof.)

"Optical and Radio Tracking of Satellites."

report presented at the Intl. Symposium on Space Age Astronomy,  
Pasadena, California, 7-9 Aug 61.

Vice Pres, Astronomical Council, Acad. Sci. USSR

S/030/61/000/004/009/015  
B105/B206

AUTHOR: Masevich, A. G., Doctor of Physical and Mathematical Sciences  
TITLE: Three-year experience in the observation of artificial earth satellites  
PERIODICAL: Vestnik Akademii nauk SSSR, no. 4, 1961, 94-98

TEXT: The international observation of artificial earth satellites is described. In 1957, a special "service for satellites" was established within the program of the International Geophysical Year, consisting of a network of observation stations and beginning to operate after the launching of the first Soviet satellite on October 4, 1957. There are 96 such stations in the USSR, 110 in the USA, 80 in Japan, 24 in China, 9 in Poland, 7 in East Germany, 5 in Czechoslovakia, 2 each in Rumania and Yugoslavia and 1 each in Bulgaria, Mongolia, and Vietnam. From January 20-25, 1961, a Conference was held at the Dom druzhby s narodami zarubezhnykh stran (House of Friendship With Foreign Peoples) which was convened by the Astronomicheskii sovet Akademii nauk SSSR (Astronomical Council of the AS USSR). This Conference was also attended by delegates from Bulgaria, Hungary, East Germany, China, Mongolia, Poland, and

Card 1/4

S/030/61/000/004/003/015  
B105/B206

Three-year experience ...

Czechoslovakia. Observations of Soviet satellites are reported to the USSR by a number of socialist and foreign stations and observatories. The method of visual observation is nearly the same in all stations. Telescopes of the types AT-1 and T3K (TZK) are mainly used. The method of visual and photographic observation as well as the possibilities of defining the data were discussed at the Conference. Reports were also made on new and perfected devices. The study of the gravitational field of the earth by means of analyzing the movements of the third Soviet satellite, conducted by I. D. Zhongolovich, which he concluded at the Institut teoreticheskoy astronomii Akademii nauk SSSR (Institute of Theoretical Astronomy AS USSR) is mentioned as an example. Investigation of the movement of artificial satellites led to the development of a new branch of gravimetry, the astronomicheskaya gravimetriya (Astronomical gravimetry). Its task consists in investigating the parameters of the potential of the gravitational force and in establishing a precise theory of the motion of cosmic bodies close to the earth. A novyy otdel prikladnoy nebesnoy mekhaniki (Department of Applied Celestial Mechanics), which is to evaluate the observations of the satellites and to calculate their orbits, was established at the Institute. The evaluation is done on the electronic computer of the type BESM (BESM) of the Vychislitel'nyy tsentr Akademii nauk SSSR (Computer Center of the Academy of Sciences

Card 2/4

S/030/61/000/004/009/015  
B105/B206

Three-year experience ...

USSR), air density in great altitudes being defined and the ellipticity of the earth determined. Accurate photographic observations must be consulted for the determination of important parameters. L. Sekhnal investigated the effect of the pressure of the earth on the motion of near satellites at the Astronomical Institute in Ondrzheyov (Czechoslovakia). At the Potsdam Observatory, East Germany, V. Günzel-Lingner made an experiment using a big long-focus double astrograph for very accurate observations of satellites. He also elaborated a simple method for an observation station to explain ephemerides. M. K. Abele of the Riga Station and L. A. Panayotov of the Pulkovo Observatory designed special cameras for photographing weakly visible satellites. I. Almar, Budapest Observatory, Hungary, developed the "Navikord" computer which simplifies the evaluation of observations. The Polish National Committee for International Geophysical Collaboration developed a special field of research, which directly evaluates observation data of satellites. The observation results of the Soviet satellites by Soviet as well as foreign stations are regularly published in special bulletins by the Astronomical Council. Data of scientific research as well as descriptions of new devices and the perfection of the method are published in the "Byulleten' stantsiy opticheskogo nablyudeniya sputnikov" (Bulletin of the

Card 3/4

Three-year experience ...

S/030/61/000/004/003/015  
B105/B206

Optical-observation Stations for Sputniks), which is also published by the  
Astronomicheskiiy sovet (Astronomical Council). Detailed recommendations  
regarding all problems dealt with were accepted by the Conference.

Card 4/4



MASEVICH, A.G., doktor fiz.-matem.nauk

Extragalactic astronomy and cosmology. Vest. AN SSSR 31 no.11:  
106-109 N '61. (MIRA 14:11)  
(Astronomy) (Cosmology)

MASEVICH, A.

Symposium "Astronomy of the Space Age." Astron.zhur. 38 no.6:  
1129-1132 N-D '61. (MIRA 14:11)  
(Astronomy--Congresses) (Space flight--Congresses)

MASEVICH, A.G.

Interpretation of the evolution of main-sequence stars. Soob.  
GAISH no.105:3-19 '61. (MIRA 14:8)  
(Cosmogony) (Stars)

MASEVICH, A.G.

Evolution of main-sequence stars of early spectral classes. Soob.  
GAISH no.105:20-34 '61. (MIRA 14:8)  
(Cosmogony) (Stars)

MASEVICH, A.G.

Nature of young stars reaching the main sequence. Soob.GAISH  
no.105:35-49 '61. (MIRA 14:8)  
(Cosmogony) (Stars)



1-22131-65

ACCESSION NR: A5001308

cession of some prospects for future developments. Author's summary.

SUB CODE: SV AA

ENCL: 00

000-1/2

MASEVICH, A.G., doktor fiz.-matemat.nauk

Advances in astronomy. General Assembly of the International Astronomical Union. Priroda 51 no.1:72-88 Ja '62. (PIRA 15:1)  
(Astronomy--Congresses)



1. ACHUTCH, A.G., doktor fiz.-matem.nauk

Mysteries of the universe are being unveiled. Priroda 52  
no.9:17-20 S '62. (MIRA 1960)  
(Astronautics)

~~MASEVICH, A.G.~~ doktor fiziko-matem. nauk, otv. red.; GHS'K'V, G.G.,  
red.izd-va; YEGOROVA, N.F., tekhn. red.

[Origin of life in the universe] Vozniknovenie zhizni vo Vse-  
lennoi; sbornik dokladov. Moskva, 1963. 95 p. (MIRA 16:9)

1. Soveshchaniye Komissii po Kosmogonii Astronomicheskogo  
soveta AN SSSR. Moscow, 1962.  
(Life on other planets) (Life—Origin) (Cosmogony)

BAKULIN, P.I., otv. red.; DAGAYEV, M.M., red.; KULAGIN, S.G.,  
red.; KUROCHKIN, N.Ye., red.; MASEVICH, A.G., red.;  
RAKHLIN, I.Ye., red.; SHKLYAR, S.Ya., tekhn. red.

[Astronomical calendar: Yearbook. varying part, 1964] Astronomi-  
cheskii kalendar'. Ezhegodnik, peremennaya chast', 1964. Red. koll.  
P.I. Bakulin i dr. Moskva, Fizmatgiz, 1963. 279 p. (Vse-  
soiuznoe astronomo-geodezicheskoe obschestvo, no. 67)  
(MIRA 17:1)

L 32667-55 PSS-2/PSE(h)/FBD/EWT(d)/EEO-2/EWT(1)/FS(v)-3/EBU(k)-2/EWA(h)/T/  
 ERI(c)-2/EBD-2/EWA(c)/EBD(b)-3 Pn-1/Po-1/Pac-1/Pq-1/Pas-2/Pg-1/Pk-1/P1-1 IJP(c)  
 ACCESSION NR: AT500474 GW/GN S/3126/63/000/002/0153/0157

AUTHOR: Zakhovoy, D. L. (Representative of PAN to Commission for multilateral cooperation); Maslov, A. G. (Representative to Commission for multilateral cooperation)

SUBJECT: Second conference of the Committee on Multilateral Collaboration on the Problem of Optical Observation of Artificial Earth Satellites, Warsaw, June 1963

SOURCE: Natsionalnyy iakustvannykh sputnikov Zemli, no. 2, 1963, Warsaw, PAN, 1963, 157-167

TOPIC: Optical satellite, satellite tracking, NAFA 38/25 camera, Echo I satellite

ABSTRACT: This conference convened simultaneously with the fourth symposium of COSPAR. Participating countries were Hungary, East Germany, Poland, USSR, and Czechoslovakia, with observers from China. It was reported that the INTERORS program, under the direction of Dr. J. L. of Hungary and set up during the first conference, was continued during the year. It was recommended that all observing stations be equipped with appropriate cameras and that synchronous photographic observations be made. Stations at Potsdam, Bucharest, Prague, Poznan, and Patany were added to the Soviet stations. Different coordinators were appointed for

Card 1/2